

1 **Title:** Superposition eyes of diurnal and nocturnal hawkmoths are imperfect spheres and lack
2 acute zones.

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21 **Abstract:**

22
23 Superposition compound eyes improve sensitivity by pooling light from multiple facets and are
24 widespread among nocturnal insects, including moths and beetles. Optical theory predicts that
25 superposition eyes must be nearly spherical to form coherent images and therefore lack highly
26 pronounced acute zones with high spatial resolution. To examine this, we imaged eyes of six
27 hawkmoths (Sphingidae) spanning diurnal and nocturnal activity with high-resolution
28 microcomputed tomography. Our automated pipeline created detailed eye maps quantifying
29 morphological parameters. We measured local eye curvature (radial distance), interommatidial
30 angle ($\Delta\phi$), facet size and crystalline cone skewness (tilt of cone axes relative to the local surface
31 normal). All species show more curvature in the dorso-ventral plane with flattening in the antero-
32 posterior plane. However, their eyes still retain near-spherical curvature globally, with diurnal
33 species showing greater distortion. For facet parameters, spatial acuity is generally highest
34 (lowest $\Delta\phi$) near the eye center and decreases gradually toward the periphery. However, overall
35 variation in spatial acuity is low and these eyes lack distinct acute zones. Facet size gradually
36 changes from center to periphery, increasing in some species and decreasing in the others. Cone
37 skewness is present in all six species (0° - 10°), but in two diurnal species of hummingbird
38 hawkmoths it increases markedly in the posterior region (15° - 30°) possibly compensating for
39 regional eye flatness. This paper provides foundational data of ommatidial and eye shape
40 measurements and advances our assumptions about how superposition eyes function.

41 **Introduction**

42
43 Hawkmoths or Sphingidae are a diverse group of robust moths with over 1700 species
44 and varied life histories (Kawahara et al. 2009; Kawahara and Barber 2015; Kitching et al.
45 2018). Their evolutionary history includes multiple transitions between diurnal and nocturnal
46 adult lifestyles within a single clade (Kawahara et al. 2009; Akiyama et al. 2022). Hawkmoths

48 are also model organisms for flight and sensory biology, able to navigate through cluttered
49 vegetation in quite dim light (Stöckl and Kelber 2019), and discriminate color at night (Kelber et
50 al. 2002; van der Kooi et al. 2021). Their visual ecology has been studied in depth for a handful
51 of nectar-feeding, hovering species from the Macroglossinae and Sphinginae subfamilies (Stöckl
52 and Kelber 2019) but not in Smerinthinae, which are predominantly non-hovering and non-
53 nectar feeding as adults (Kawahara et al. 2009; Kawahara and Barber 2015; Kitching et al.
54 2018). The visual systems of nectar feeding and hovering hawkmoths have likely been optimized
55 for three-dimensional flight control and rapid forward flight in their preferred light environment
56 (Stöckl et al. 2017a). Acute and sensitive vision supports precision flight by stabilizing scenes
57 during hovering and processing rapid optic flow during forward motion (Sponberg et al. 2015;
58 Stöckl et al. 2017a; Menz et al. 2022).

59 Hawkmoths have superposition compound eyes. These eyes differ markedly from single
60 lens camera-type eyes, such as those found in vertebrates or spiders. They are made up of arrays
61 of visual units, called ommatidia, that are typically packed in a hexagonal lattice and can be
62 categorized into two kinds, superposition or apposition (Stavenga 2006). Most day-active insects,
63 such as bees and butterflies, possess **apposition** compound eyes, in which each facet focuses
64 light onto a fused bundle of photoreceptor microvilli, the rhabdom (Land and Nilsson 2012).
65 Light along a facet's anatomical axis reaches the rhabdom, but off-axis light is absorbed by
66 surrounding screening pigments. One can think of a rhabdom as a pixel (although each rhabdom
67 can contain microvilli from 5–9 photoreceptors) with each facet directing light onto a single spot
68 in an image (Pichaud and Casares 2022). Insects active in dim light, such as moths, beetles, and
69 owlflies, instead possess **superposition** compound eyes, an adaptation that greatly enhances
70 light-gathering ability and sensitivity affording better night vision (Warrant and McIntyre 1990;
71 Warrant 2001; Meyer-Rochow and Gál 2004; Belušič et al. 2013). Unlike apposition eyes, which
72 screen out off-axis light, superposition eyes have a clear zone between lenses and retina, and
73 optics that redirect incoming light from many facets onto a single rhabdom. Optical theory
74 suggests that the increase in sensitivity comes with trade-offs, one of them being that
75 superposition eyes need to be spherical and have equally sized facets across the eye to
76 successfully implement superposition (Warrant et al. 1999; Stavenga 2006)

77 Moths and butterflies have transitioned between diurnal and nocturnal lifestyles many
78 times in their evolution (Kawahara et al. 2018) and correspondingly they have both apposition
79 and superposition eyes and some intermediate forms (Fischer et al. 2014). The distribution and
80 evolution of superposition and apposition eyes across Lepidoptera is unclear, but studies have
81 proposed links to preferred activity times or diel niche (Horridge et al. 1977; Land 1984; Greiner
82 2006; Yack et al. 2007; Fischer et al. 2012). However, because diel niche often represents a
83 spectrum rather than a distinct set of categories and can also change in a species (Kirse et al.
84 2025), disentangling eye and diel-evolution can be difficult.

85 To understand the relationship between superposition architecture and diel niche, we
86 examined the eyes of six sphingid moth species, three with fully or partial diurnal activity and
87 three with mostly nocturnal activity. We used microcomputed tomography (μ CT) to reconstruct
88 the internal and external eye structure, then examined compound eye parameters as they varied
89 across the eye, either from center to periphery or dorsal to ventral, using specialized software
90 (Currea et al. 2023). We estimated local sphericity using radial distance, eye size, facet size, cone
91 skewness and interommatidial angles ($\Delta\phi$). Crystalline cones are not always perfectly
92 perpendicular to their facet surface; some are skewed, causing the ommatidium to view along a
93 slightly oblique axis. We used ODA to first estimate the anatomical axes of each cone. This

94 allowed us to estimate (1) cone skewness as the angle between a tangent from the eye surface
95 and the cone axis and (2) the $\Delta\phi$ as the mean visual angle between each cone and its six
96 neighbors (Curra et al. 2023). We tested the idea that superposition eyes need to be spherical
97 and examined if diurnal and nocturnal moths both held to this constraint (Stavenga 1979, 2006;
98 Warrant et al. 1999). We also tested if superposition eyes would lack regional specialization or
99 acute zones and examined how acuity, facet size and skewness changed across the eye.

100

101 **Methods**

102 *Sampling design and choice of study species*

103 We chose six hawkmoth species from two subfamilies (Macroglossinae and Sphinginae)
104 representing different degrees of diurnality and nocturnality as adults: one species is exclusively
105 diurnal (*Macroglossum stellatarum*), two species are both diurnal and nocturnal (*Macroglossum*
106 *fritzei*, *Hyles lineata*), and three species are primarily nocturnal (*Deilephila elpenor*, *Sphinx*
107 *ligustri*, *Manduca sexta*). Diurnality and nocturnality designations were determined based on
108 published literature and consultation with experts. The species selected, along with information
109 on the specimens used, are listed in Supplemental Table 1.

110

111 *Microcomputed tomography*

112 To visualize internal eye structure, we conducted μ CT scanning at both the Natural
113 History Museum (NHM), London, UK and the University of Florida (UF), Nanoscale Research
114 Facility, Gainesville, FL, USA. Below we outline the protocol followed at each institution.

115 For specimens imaged at UF, we collected the heads of each species and processed them
116 with a standardized preparation protocol for μ CT scanning (Bray et al. 1993; Metscher 2009).
117 Each specimen was initially preserved in a 1.5 ml Eppendorf tube containing 70% ethanol for at
118 least 24 hours. Following preservation, we replaced the ethanol with a 0.3% phosphotungstic
119 acid (PTA) solution prepared in 70% ethanol and allowed specimens to stain for at least three
120 weeks. In most cases, we scanned both eyes, but as some specimens showed collapse on either
121 side, we sometimes only scanned a single eye.

122 Eyes larger than 3 mm in diameter were removed directly from PTA and scanned in a wet
123 state, (ZEISS Versa 620 XRM). For eyes smaller than 3 mm, dehydration during scanning
124 caused significant shrinkage and displacement, leading to motion blurring. To prevent this, we
125 scanned smaller eyes in a dry state.

126 To prepare dry samples, we followed a graded ethanol dehydration process. Specimens
127 were sequentially immersed in 80%, 90%, and 100% ethanol, each for 1 hour. Following
128 dehydration, samples were transferred to a 1.5 ml Eppendorf tube containing
129 hexamethyldisilazane (HMDS) for 24 hours, to facilitate infiltration and stabilize internal eye
130 structures. After infiltration, samples were removed from the HMDS solution and placed in a
131 dry, capped Petri dish to allow for slow evaporation. All μ CT scans at the University of Florida's
132 Nanoscale Research Facility were performed using Phoenix vjtomex M (GE Measurement &
133 Control Solutions) or a ZEISS Xradia 620 Versa 3D X-ray microscope. The ZEISS system is
134 equipped with an automatic reconstruction feature that processes scan data immediately after a
135 complete scan.

136 For specimens scanned at the NHM, mouthparts and antennae were removed from each
137 head using a dissection microscope to increase the staining efficiency of internal eye structures.
138 The specimens were then fixed in 70% ethanol for at least 24 hours. Following fixation, the
139 ethanol was replaced with 1% PTA solution prepared in 70% ethanol and then allowed to stain

140 for seven days. To examine the effectiveness of staining, specimens were individually scanned.
141 When staining was insufficient, we placed the specimen back into 1% PTA solution and scanned
142 every two days until staining was complete, or for a maximum of 21 days. These μ CT scans
143 were performed using a ZEISS Xradia 520 Versa 3D X-ray microscope.

144 145 *Scan data processing*

146 After μ CT reconstruction, we aligned each scan to a consistent dorso-ventral and
147 anterior-posterior orientation in 3D Slicer (Slicer 5.6.2; Fedorov et al. 2012). We manually
148 extracted ommatidial traces using segmentation tools in 3D Slicer for one eye on each specimen.
149 Segmentations were exported as NRRD files and converted to TIFF stacks via the pynrrd
150 package (See supplementary file 3 for custom code). Raw volumes are available on
151 MorphoSource, and manually traced datasets are on Figshare
152 (<https://figshare.com/s/44be67162143d41fddb6>).

153 154 *Automated measuring of ommatidia with the ommatidia detecting algorithm (ODA)*

155 We automated the analysis of pre-filtered stacks using an established ODA pipeline
156 (Currea et al., 2023). The ODA applies a fast Fourier-transform-based filter to detect the
157 hexagonal lattice of ommatidia in each 2D slice, automatically identifying individual facets and
158 computing their centroid coordinates. For every CT scan, the pipeline outputs two files
159 (ommatidial_data.csv and interommatidial_data.csv, see Supplemental data) with ommatidial
160 counts, center locations, skewness, diameter and radial distances of each crystalline cone. We
161 measured the $\Delta\phi$ of each pair of adjacent ommatidia (Supplemental File:
162 interommatidial_data.csv) and the mean $\Delta\phi$ of each ommatidium and its six neighboring
163 ommatidia (Supplemental File 2: ommatidial_data.csv).

164 165 *Defining and measuring eye parameters*

166 *Radial distance:* This is defined as the distance from the center of the best-fitting sphere
167 to the lens centroids fit using a least squares algorithm based on (Jekel et al. 2016) and further
168 described in (Currea et al. 2022).

169 *Sphericity:* We calculated two putative measures of sphericity. A) $1 - (\text{std. dev}(r) /$
170 $\text{mean}(r))$, where r is the set of all radial distances. This is a complement of the coefficient of
171 variation of eye radial distance across all ommatidia, such that as the standard deviation of radii
172 (normalized by mean radius) approaches 0, this value approaches 1 as it would for a perfect
173 sphere. B) Inner radius/outer radius: we fit two spheres and compare the radius of the minimum
174 bounding and maximum inscribing spheres (Fig. 1).

175 $\Delta\phi$: We calculated $\Delta\phi$ as the mean angular separation between the visual axis of a
176 crystalline cone and those of its six nearest neighbors.

177 *Lens diameter:* The lens diameter of a given ommatidium was approximated by its mean
178 center-to-center distance between adjacent ommatidia. The lens area of a given ommatidium was
179 approximated by assuming a circular aperture with a diameter equal to the lens diameter. We
180 used the mean radial distance as a measure of eye size. For each species, we altered the
181 algorithm parameters, such as density range and window size, to let the program recognize every
182 ommatidium (see code Supplementary file 3).

183 *Cone skewness:* Skewness for each ommatidia is the tilt of cone axes, relative to the local
184 surface normal. It was calculated as the angular deviation of the cone axis from the local facet

185 normal, which is equivalent to 90° minus the angle to the tangent plane (Stavenga 2006).

186

187

188 *Manual measuring of ommatidia*

189 To validate some of the automated measurements, we manually measured the facet or lens
190 diameter and ($\Delta\phi$) along three axes passing through the center of the eye: equatorial, 120°
191 above, and 120° below the equator. Along each axis, we sampled five regions, each spaced by
192 $\sim 15^\circ$ of viewing angle. We made the measurements in 3D Slicer on the 2D cross-section panel
193 oriented to the tested axis. In each sampled region, we measured the lens diameter by taking the
194 average end to end width of three successive facets. For the $\Delta\phi$, we traced back 5 ommatidia to
195 the point of intersection and divided the angle by 5. We also estimated the eye size of each
196 species by measuring the diameter of the eye horizontally.

197

198 *Capturing representative 2D images*

199 After reorienting each scan in 3D Slicer, we generated two orthogonal cross-sections
200 passing through the estimated geometric center of each eye. One section was in the transverse
201 plane (horizontal to the eye) and the other in the sagittal plane (vertical to the eye).

202

203 *Statistical analyses and measuring eye parameters*

204 We calculated the mean and 99.9% confidence intervals of the mean for lens area, $\Delta\phi$,
205 skewness angle, and radial distance of the facets and then measured the Pearson correlation of
206 lens area, $\Delta\phi$, and cone skewness relative to mean eye radius. Parameter definitions and
207 computational details follow Currea et al. (2023); briefly, we measured lens diameter, lens area,
208 $\Delta\phi$, cone skewness, and radius in both transverse and sagittal planes. Using ordinary least
209 squares regression, we modeled lens area, which limits optical sensitivity, as an affine function
210 of skewness angle.

211

212 **Results**

213 We successfully constructed μ CT scans of the six hawkmoth species (Fig. 1). We then
214 oriented all reconstructions in the same direction, cleaned and cropped each to a single eye, and
215 manually cleaned until the crystalline cones became visible (Fig. 2, S1, S6). We then estimated
216 all eye parameters with the ODA and manually aligned the final maps with the CT scans to
217 identify the dorso-ventral and antero-posterior axes (Fig. 3)

218

219 *Ommatidial counts, facet size, $\Delta\phi$ and correlations with eye size*

220 We first quantified the number of ommatidia and variation in lens area across species. Counts
221 ranged from 5,500 to 28,800 ommatidia (Table 1, Fig 1A) and increased with eye diameter ($r =$
222 0.95 , $p = 0.036$; Fig. 1B). The diurnal species *Macroglossum stellatarum* had the smallest eyes
223 and lowest ommatidial counts ($\sim 5,540$), while the crepuscular *Manduca sexta* had the largest
224 eyes and highest counts ($\sim 28,800$). Facet size (lens area) was positively correlated with eye size
225 ($r = 0.86$, $p = 0.029$, Fig. 1B), with the smallest in *Macroglossum stellatarum* and largest in
226 *Manduca sexta*. $\Delta\phi$ was negatively correlated with eye size ($r = -0.69$, $p = 0.13$, Fig 1B), with
227 *Manduca sexta* exhibiting the lowest $\Delta\phi$ (highest spatial acuity).

228

229 *Eyes are close to spherical, but have consistent local changes in radial distance*

230 We measured eye size using automated fitting of ommatidial centers to a spherical surface (Table
231 1) and manual measurements of eye diameter (Table S2). Radial distance varied systematically
232 across the eye surface (Fig. 3), revealing consistent local departures from a perfect sphere. We
233 found that in all six species, the eye was flatter along the antero-posterior axes and more curved
234 along the dorso-ventral axis (Fig 3, S7). We estimated global sphericity using the coefficient of
235 variation of radial distance and found that all species eyes had values close to 1, but when fitting
236 a minimum bounding and maximum inscribing sphere (inner/outer radius) the diurnal species
237 showed more deviations from a perfect sphere than nocturnal species (Fig 1B, See methods for
238 more details).

239

240 ***$\Delta\phi$ increases from center to periphery in most nocturnal species***

241 $\Delta\phi$ varied systematically across the eye axes (Figs. 4, S2, S3). In all species, $\Delta\phi$ was lowest in
242 the central region of the eye and increased towards the periphery (Fig. 4, Table S2). Local
243 irregularities in $\Delta\phi$ values were visible across the eye surface in some species (Fig. 4), but these
244 did not persist in the binned means (Fig. 4, S3). Among the six species, *Macroglossum fritzei*
245 showed the most abrupt shifts in $\Delta\phi$, whereas the remaining species showed more gradual
246 changes.

247

248 ***Facet size increases and decreases gradually from center to periphery in different species***

249 Lens diameter and area varied systematically across eye axes, but the pattern was not consistent
250 across species (Figs. 1B, Fig 5, S4, S5). We collapsed variation onto a single peripheral ring
251 (eccentricity from 0–90°, Fig. 5) and found that some species show clearer trends: *Manduca*
252 *sexta* exhibited a ~15–18% decrease in facet diameter from center to periphery, whereas both
253 *Macroglossum* species showed steeper declines (~20–23%). In all tested nocturnal species, we
254 observed only gradual decrease from center to periphery along both the dorsoventral and
255 anteroposterior axes (Fig. S5) confirmed with manual measurements (Table S2). Notably,
256 *Macroglossum stellatarum* and *Macroglossum fritzei* had a distinct large facet-size zone in the
257 dorsoposterior region, clearly visible in the eye maps and the binned means (Figs. S4).

258

259 ***Nocturnal species show minimal cone skew***

260 All three nocturnal species exhibited minimal skew across both axes (Fig. 6). In contrast,
261 *Macroglossum fritzei* and *Macroglossum stellatarum* displayed a sharp increase in skew toward
262 both the posterior and anterior regions of the eye, exceeding 20 degrees in the posterior of *Mac.*
263 *fritzei* (Fig 1B). *Hyles lineata* showed a similar but smaller increase towards the posterior eye.
264 Using ordinary least squares regression, we modeled lens area, as an affine function of skewness
265 angle. We found significant positive correlations between skew angle and lens area in the 4
266 *Macroglossiinae* ($r > 0.18$, $p < 1 \times 10^{-57}$) but not in the 2 *Sphinginae* ($r < 0.002$, $p > 0.73$; Table
267 S3). Lens area increased by 4.1 to 4.8 μm^2 for every degree increase in skewness angle, as
268 opposed to 0.02 to 0.03 $\mu\text{m}^2/\text{deg}$ in *S. ligustri* and *M. sexta*.

269

270 **Discussion**

271 ***Optical theory of superposition eyes***

272 In a dark-adapted superposition eye, each rhabdom receives light from many corneal facets,
273 which greatly increases sensitivity. But this requires that each facet, in turn, correctly directs
274 light towards many rhabdoms, depending on angle of incidence, and this significantly constrains
275 the eye's geometry. This implies superposition eyes might lack the regional specializations that

276 characterize many apposition eyes, as the surface may need to remain nearly spherical, with
277 facets of relatively uniform spacing and size. Our results generally support this view. Across six
278 sphingids displaying both diurnal and nocturnal activity, we found their eyes remaining close to
279 spherical. $\Delta\phi$ varied gradually rather than forming discrete acute zones, and facet sizes changed
280 smoothly across the eye surface. However, we did find repeatable deviations from perfect
281 spherical symmetry. All tested species showed flattening in the antero–posterior plane and
282 greater curvature in the dorso–ventral plane, indicating that superposition optics tolerate modest
283 regional distortion. Further, diurnal species showed stronger local departures, including
284 pronounced crystalline cone skew in posterior regions of *Macroglossum*. This suggests
285 superposition eyes are not rigidly constrained to perfect spherical geometry but operate within an
286 envelope which can compensate for moderate distortions—potentially through cone skew or
287 subtle changes in facet geometry. So rather than prohibiting regional specialization,
288 superposition architecture may simply limit its magnitude and sharpness.

289

290 Diel-evolution and previous visual ecology studies in hawkmoths.

291 Although most hawkmoths are night-flying, some are strictly diurnal and others show
292 both day and night activity (Beck and Linsenmair 2006; Broadhead et al. 2017). Within
293 hawkmoths, diurnal reversions have occurred independently in several genera, sometimes
294 involving only a small number of species across three of the four subfamilies. These include
295 *Afrosatspes* and *Satsapes* (subfamily Smerinthinae), one species of *Sphinx* (subfamily
296 Sphinginae), and *Cephonodes* and *Hemaris*, *Aellopos*, *Euproserpinus* and most *Proserpinus*, as
297 well as *Hayesiana* and some species of *Macroglossum* (subfamily Macroglossinae). However,
298 within these diurnal species, vision has only been studied extensively in the hummingbird
299 hawkmoth, *Macroglossum stellatarum*. In this species, the superposition pupil (the effective
300 superposition aperture visible as the eye glow) remains unusually open even in bright light, so
301 the eye retains a refracting superposition state under conditions in which many moths would
302 reduce optical pooling through pigment migration (Warrant et al. 1999; Stöckl et al. 2016, 2022).
303 Relative to more nocturnal hawkmoths, however, the superposition aperture is smaller, consistent
304 with adaptation to brighter visual environments (Stöckl et al. 2017b). No other diurnal
305 hawkmoths superposition architecture has been studied in similar detail, making it difficult to
306 determine whether these features are exceptional or a pattern present across all diurnal
307 hawkmoths. To explore how eye architecture relates to diel niche, we therefore selected a mix of
308 diurnal and nocturnal species, prioritizing species with known visual ecology, or those with
309 accessible colonies, and explored using 3D eye scans.

310

311 Hawkmoth eyes are near-spherical in shape, with axis-specific departures

312 Quantifying sphericity in biological eyes is inherently metric-dependent, because
313 “sphericity” is not uniquely captured by any single index. We therefore report two
314 complementary metrics (globally and locally) as descriptive summaries of eye-shape variation.
315 Results from both metrics suggest minimal deviation from a spherical eye shape. *Macroglossum*
316 *stellatarum* exhibits the lowest sphericity among the species examined, but still closely
317 approximates a sphere—especially compared to the much less spherical eyes in many other
318 insects.

319 Within each eye, we found consistent variation in estimated curvature radius in both eye
320 axes dorso-ventrally and antero-posteriorly across all species, providing a measure of how flat or
321 curved the eye is at different regions. This suggests that all species show regions of local

322 flattening and regions with more curvature. Interestingly, the orientation of this curvature
323 variation was not consistent across species, not even between the different *Macroglossum*
324 species. This suggests that species may be optimizing the balance between field of view (more
325 curved) and higher acuity (flatter), in ways that match their flight and motion requirements. This
326 warrants further exploration in future studies, as it could reflect a developmental constraint as
327 much as an adaptive benefit.

328

329 Nocturnal species have larger eyes featuring larger and more numerous facets than the strictly
330 diurnal species

331 Our results show that strictly diurnal species have the smallest eyes, with the fewest
332 ommatidia and smallest facet size (Table 1; Fig. 1B). Across the six species, eye size was
333 positively correlated with both facet count and facet size (Fig. 1B), consistent with the
334 expectation that larger eyes accommodate more sampling units and larger optical apertures. We
335 did not find evidence that strictly nocturnal species necessarily have the largest eyes in this
336 dataset, because *Hyles lineata*, which is active both during the day and at night, had the second-
337 largest eye size and ommatidial count. Our results are consistent with results published by (Yagi
338 and Koyama 1963), which reported similar scaling between eye size, facet number, and facet
339 dimensions in Lepidoptera.

340 In insects with apposition eyes, larger facets in nocturnal taxa are commonly interpreted
341 as an adaptation for increased photon capture and thus higher optical sensitivity in dim light
342 (Somanathan et al. 2009). Although superposition eyes differ in that sensitivity is strongly
343 shaped by optical pooling and by pupil dynamics, facet size and eye size should still contribute
344 to potential light capture when the superposition pupil is open. We did not quantify pupil size
345 under fully dark-adapted conditions in our specimens, so we cannot estimate maximal optical
346 sensitivity directly from our scans. Nevertheless, the fact that nocturnal species retain larger
347 facets and larger eyes even in the light-adapted state suggests that these differences are not solely
348 a transient consequence of pupil state. Our results are consistent with the interpretation in
349 previous studies that nocturnal hawkmoths achieve higher potential sensitivity through a
350 combination of larger optical apertures and a larger pupil. (Yagi and Koyama, 1963; Stöckl et al.
351 2017b)

352

353 Intraocular variations in facet size and $\Delta\phi$

354 Both facet size and $\Delta\phi$ vary across the eye surface within each species, and these
355 intraocular gradients tend to be steeper in the diurnal species. We previously hypothesized that
356 diurnal taxa might show more abrupt regional changes in ommatidial size and spacing, and the
357 strictly diurnal *Macroglossum stellatarum* provides some support for this idea: despite having the
358 lowest facet counts, it shows the sharpest within-eye transitions in both facet size and $\Delta\phi$,
359 consistent with previous reports (Warrant et al. 1999; Stöckl et al. 2022). However, we did not
360 find evidence that nocturnal species lack regional variation across the eye; instead, they also
361 show regional change, but typically with more gradual transitions. One possible explanation is
362 that diurnal hawkmoths may operate their superposition eyes under light-adapted conditions
363 more frequently, when effective optical pooling is reduced and regional differences in
364 ommatidial size and spacing may have stronger functional consequences. This idea remains
365 speculative and will require future tests linking pupil state, light adaptation, and spatial variation
366 in optical performance.

367 Patterns of facet-size variation differed among species. *Manduca sexta* showed an
368 antero-ventral region with larger lenses that decreased in size both ventrally and toward the
369 periphery. *Sphinx ligustri* and *Deilephila elpenor* showed comparatively the least changes in
370 facet size across the eye. *Hyles lineata* showed an entirely different pattern with an almost
371 anterior dorsal bump in facet size. Taken together, these results indicate that both diurnal and
372 nocturnal taxa can exhibit pronounced regional tuning of facet size, but the form of that tuning is
373 species-specific rather than neatly partitioned by diel niche.

374 In contrast to facet size, $\Delta\phi$ was relatively uniform across the eye surface in all species,
375 with most species showing a modest increase toward the periphery (Fig 8-9). Despite the center
376 of the eye generally providing the highest spatial resolution (lowest $\Delta\phi$), none of the six species
377 exhibited a distinct acute zone in the sense of a sharply bounded region with strongly reduced $\Delta\phi$.
378 Instead, acuity changes were gradual, producing a broad central region of slightly enhanced
379 resolution rather than a pronounced fovea-like specialization. This pattern is expected because
380 coherent imaging in superposition eyes depends on maintaining near-spherical curvature, and
381 therefore major local departures such as localized flattening that would strongly compress $\Delta\phi$ are
382 limited. Taken together, our facet size and $\Delta\phi$ maps and curvature measurements support the
383 hypothesis that superposition architecture favors relatively uniform sampling with only modest,
384 gradual regional tuning, rather than highly localized acute zones.

385
386 *Cone skewness as a regional specialization in tribe Macroglossinae and genus Macroglossum*

387 Compared with the other species, *Macroglossum stellatarum* and *Macroglossum fritzei*
388 showed a pronounced posterior region of highly skewed crystalline cones, consistent with
389 previous reports from the former species (Warrant et al. 1999). This pattern is only evident in
390 antero-posterior views but not in dorso-ventral sections, suggesting that the skewness is
391 directionally structured rather than a global tilt expressed across all axes. A more plausible
392 geometric interpretation is that the pronounced posterior skewness redistributes ommatidial
393 viewing directions and likely expands posterior field of view. This tradeoff would be expected to
394 increase local $\Delta\phi$ along the antero-posterior axis and therefore reduce both acuity and sensitivity
395 in the affected region. *Macroglossum* species are exceptionally fast and maneuverable, and a
396 regional visual specialization of this kind could support accurate flower tracking with frontal
397 vision while enabling rapid responses to peripheral motion during evasive flight. No other
398 species in our sample shows comparably strong posterior skewness, suggesting this may be a
399 *Macroglossum* associated adaptation rather than a general feature of Sphingidae superposition
400 eyes

401 Cone skewness is also expected to involve optical tradeoffs. Changing cone axis
402 orientation can redistribute viewing directions and potentially expand regional field of view
403 coverage, but it can reduce sensitivity by decreasing the effective aperture and optical gain of the
404 affected ommatidia. In this context, pronounced cone skewness may be most feasible in taxa that
405 routinely operate under bright or intermediate light levels, where photon limitation is weak and
406 some loss of sensitivity can be tolerated in exchange for altered visual sampling. This provides a
407 possible explanation for why strong skewness is expressed in the diurnal *Macroglossum* species
408 but is absent from the strictly nocturnal species in this study.

409 Interestingly, we saw a phylogenetic constraint in patterns of cone skewness and lens area, with
410 significant positive correlations between skew angle and lens area in the 4 Macroglossinae (a
411 weak correlation ~ 0.2) but completely missing in the 2 Sphinginae (Table S3). No other metric
412 showed any clear phylogenetic patterns, if anything eye parameters seem to switch rapidly even

413 within a genus, so exploring evolutionary variation of skewness across hawkmoths is good
414 avenue to explore.

415 Together, these patterns suggest that, within the global geometric constraints of
416 superposition optics, regional tuning may be achieved through coordinated variation in cone
417 orientation and lens size, rather than through a sharply bounded acute zone.

418

419 *Patterns of facet and angle distribution: Lepidoptera and other insects*

420 The two central ideas we wanted to test were constraints on variation in superposition eyes and
421 how eye shape and curvature might change. To our knowledge, eye curvature has been examined
422 only by del Portillo (1936), who used histological sections from a few species and found that
423 largely the superposition eyes had equal radii of curvature in both horizontal and dorsal planes
424 and gradual changes in $\Delta\phi$ across the eyes. He examined a large Saturniidae, *Hyalophora*
425 *cecropia* (then known as *Samia cecropia* or the Cecropia silkworm), a medium sized Arctiinae,
426 *Euchelia jacobea* (then known as *Tyria jacobaeae* or the cinnabar moth) and a medium sized
427 Notodontidae moth, *Phalera bucephala* (the buff-tip moth) (del Portillo 1936). He found that the
428 *Hyalophora cecropia* eye is equally curved in both median planes, and had its curvature radii
429 homogeneous at all points, except for a small dorsal apex. The same holds for the angular values
430 (presumably $\Delta\phi$), which are virtually equivalent across most regions; except at the upper end the
431 frontal section has slightly more wide-angled ommatidia that are inclined toward the corneal
432 surface. In *Phalera bucephala*, the cross section shows less perfect regularity than in
433 *Hyalophora*, but the overall patterns are the same in both species: angular values ($\Delta\phi$) coincide
434 in the eye's center and increase slightly forward and downward. In *Euchelia jacobaeae*, the eyes
435 are small and nearly spherical, with highly uniform angles and radii; but only the upward-
436 directed ommatidia have slightly wider angles (del Portillo 1936).

437 Yagi and Koyama (1963) conducted a large-scale survey of facet variation across more
438 than 30 butterfly and moths species and described results similar to ours – a gradual drop in
439 ommatidial size from center to periphery, as well as a positive correlation between facet size and
440 eye size (Yagi and Koyama 1963). They also present facet count estimates for several
441 Sphingidae, including an estimate of 27,000 for *Agrius convolvuli* (Yagi and Koyama 1963).
442 Research on the highly sexually dimorphic moth species *Orgyia antiqua* (Erebidae:
443 Lymantriinae) revealed that females have a lower resolution ($\Delta\phi = 3^\circ$ in females, versus 1.9° in
444 males) and less curved eyes (female = $327 \mu\text{m}$ and male = $461 \mu\text{m}$), which is likely due
445 to the fact that females are flightless (Lau and Meyer-Rochow 2007). Warrant et al. (1999)
446 explored eye and facet parameter variation in *Macroglossum stellatarum* in great depth.
447 However, they concluded that optical pooling and rhabdom packing density are more responsible
448 for resolution changes than variation in facet size alone. They did find a posterior region with
449 more skewed cones and suggested this arrangement may maintain resolution needs in the frontal
450 eye, but increase field of view in the posterior (Warrant et al. 1999). Recent work using μCT
451 scans has further examined allometry and individual variation (Stöckl et al. 2022). Other moth
452 eye research compared the angular spread of eyeshine with electrophysiologically measured
453 acceptance angles, demonstrating that moths can vary considerably in whether their visual
454 systems are over or under sampling (Horridge et al. 1977). There are several groups who have
455 studied this in butterflies, such as Pieridae (Ribi 1979) and Nymphalidae (Rutowski et al. 2009),
456 but since these are largely diurnal and have apposition eyes, more facet variation is expected
457 (Seymoure et al. 2015).

458

459 Conclusion

460
461 Using high-resolution μ CT and automated whole-eye mapping across six hawkmoth species, we
462 show that hawkmoth superposition eyes are neither perfect spheres nor highly regionalized in the
463 way often seen in apposition eyes. Instead, the eyes remained broadly near-spherical and lacked
464 sharp acute zones, yet still exhibited modest regional tuning in curvature, facet size, and cone
465 orientation, especially in diurnal *Macroglossum*. Although our dataset is limited by low within-
466 species sampling and uses morphological inference alone, it provides a comparative foundation
467 for understanding how superposition eyes evolve across diel niches and for testing how local eye
468 geometry shapes visual function.

469
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479
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482 Writing – Original Draft

483 **Ruchao Qian:** Conceptualization, Investigation, Methodology, Formal Analysis, Investigation,
484 Writing – Original Draft

485 **Isabella Koushiar:** Data curation, Writing – Review & Editing

486 **Ian Kitching:** Conceptualization, Funding Acquisition, Supervision, Resources, Writing –
487 Review & Editing

488 **John Pablo Currea:** Conceptualization, Software, Formal Analysis, Writing – Review &
489 Editing

490 **Deborah Glass:** Conceptualization, Investigation, Data Curation, Resources, Writing – Original
491 Draft

492 **Jacqueline Degen:** Resources, Writing – Review & Editing

493 **Ed Stanley:** Investigation, Data Curation, Resources

494 **Simon Sponberg:** Supervision, Funding Acquisition, Writing – Review & Editing

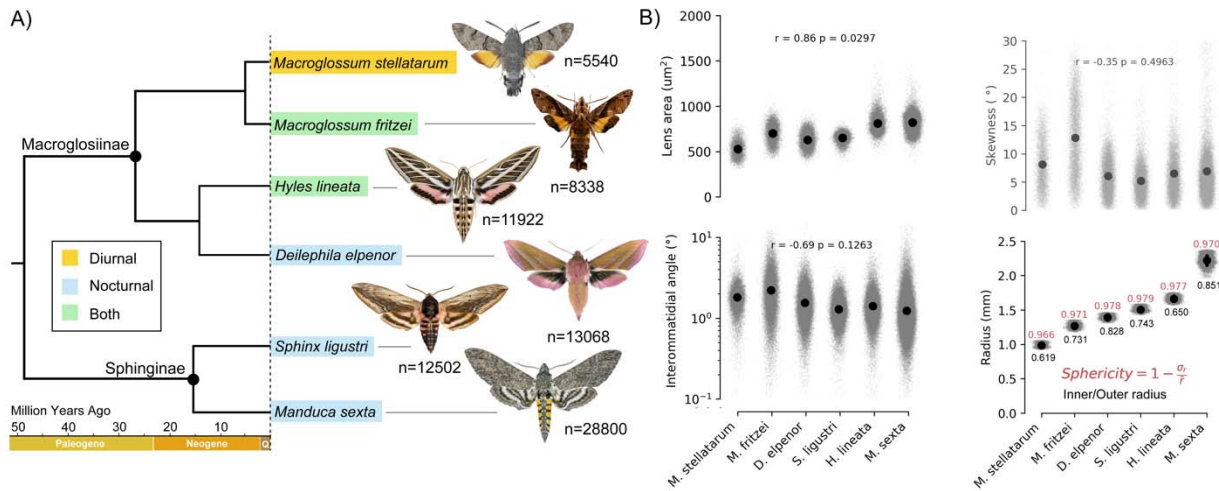
495 **Akito Kawahara:** Conceptualization, Funding Acquisition, Supervision, Resources, Writing –
496 Review & Editing

497 **Jamie Theobald:** Funding, Conceptualization, Funding Acquisition, Supervision, Resources,
498 Writing – Review & Editing

499
500 **Funding**

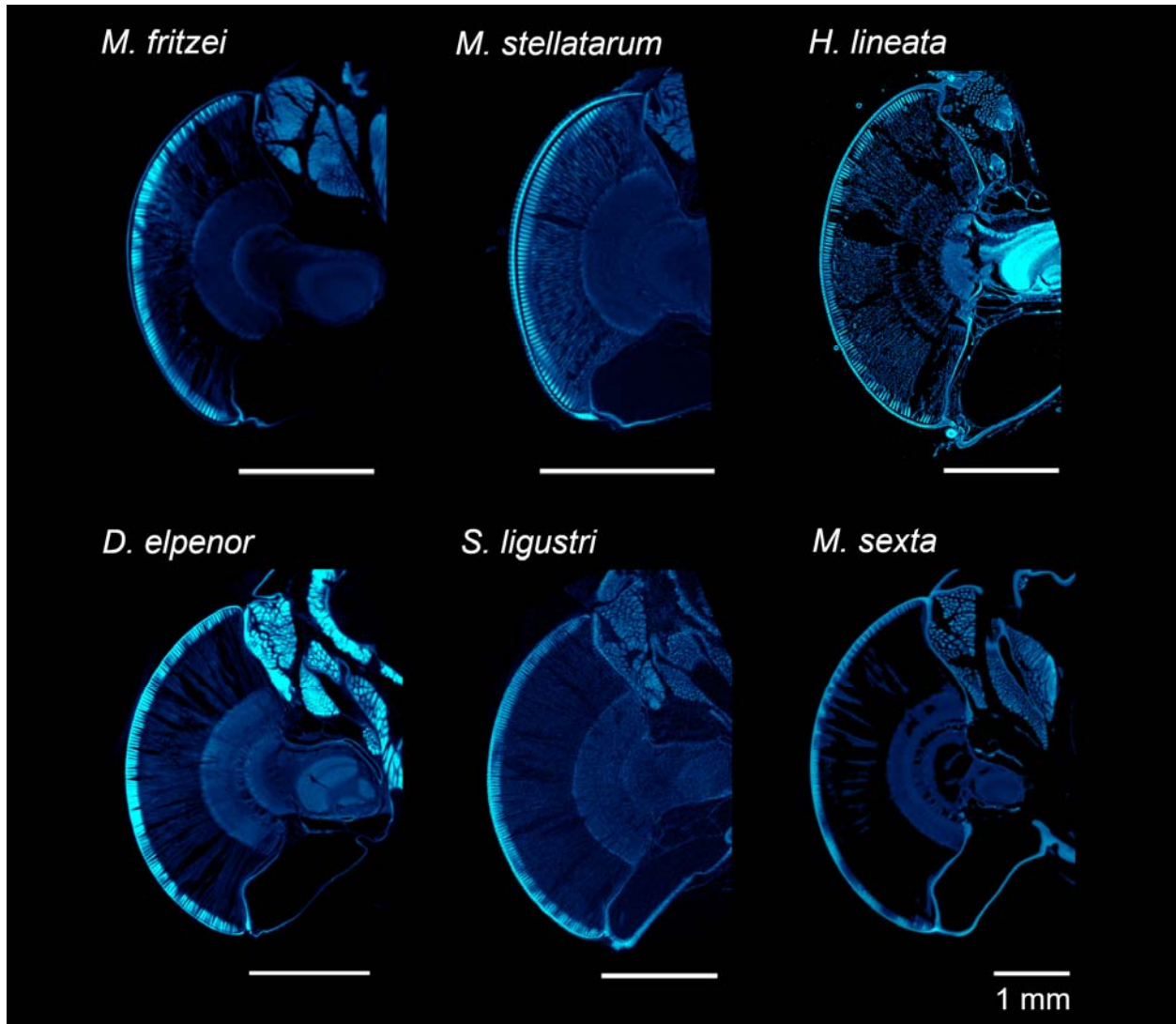
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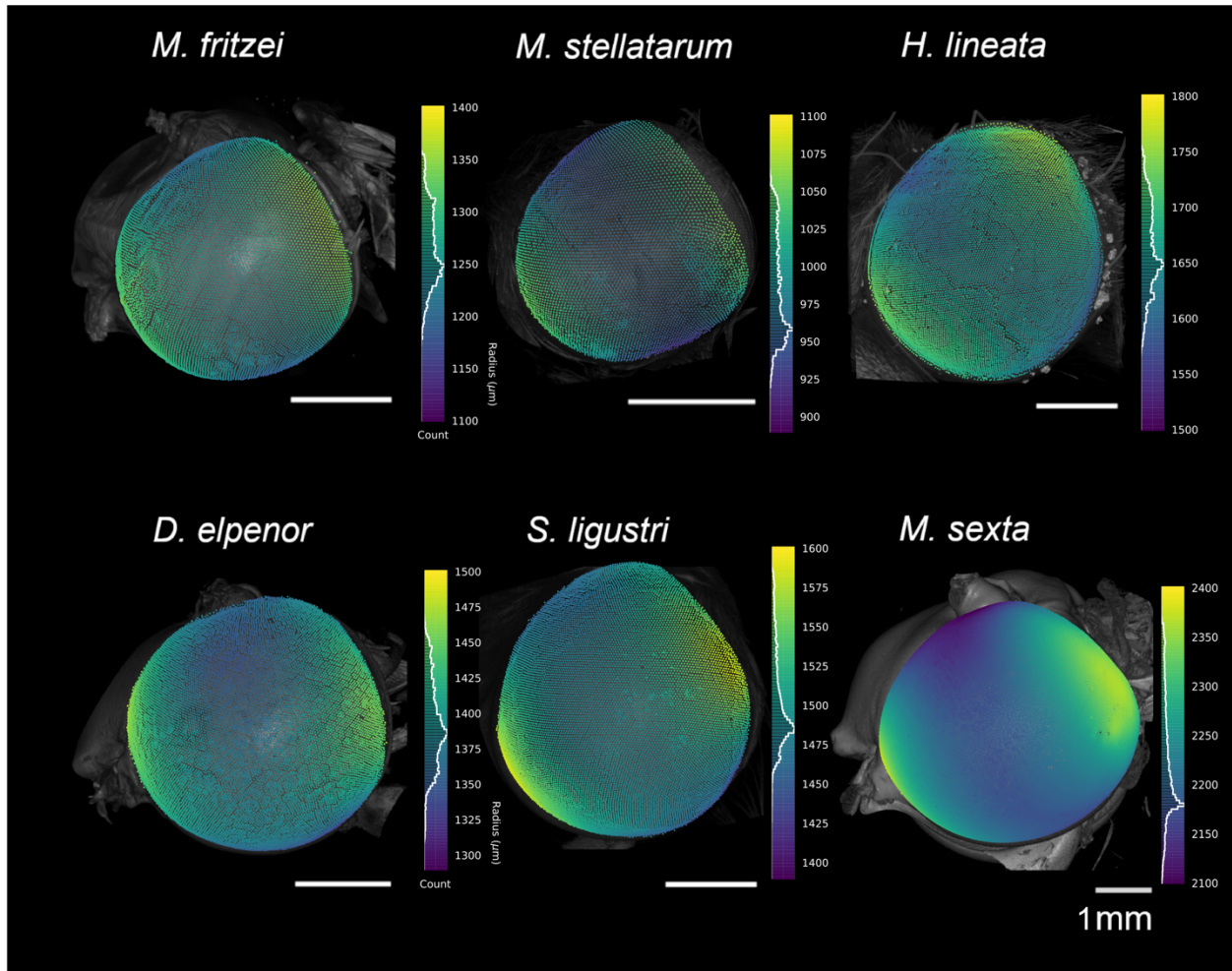
506

507 **Figure 1:** The six species tested for overall ommatidial parameters and eye sizes. A)
 508 Phylogenetic tree reconstructed from Kawahara and Barber (2015) and Kawahara et al. (2009).
 509 B) Summary statistics for the six species. Small gray dots indicate measurements for individual
 510 facets and larger black spots indicate mean values for lens area, IOA, skewness angle, and facet
 511 radius along with 99.9% confidence intervals, although these are often so small, they are hidden
 512 by the mean markers. Moth species are sorted in ascending order of mean facet radius as an
 513 estimate of general eye size. Sphericity—a measure of how close the distribution of radiuses
 514 approximates a sphere with 1 indicating a perfect sphere—is printed in red above each
 515 distribution of facet radii per species. In black below is a similar metric, inner/outer radius.

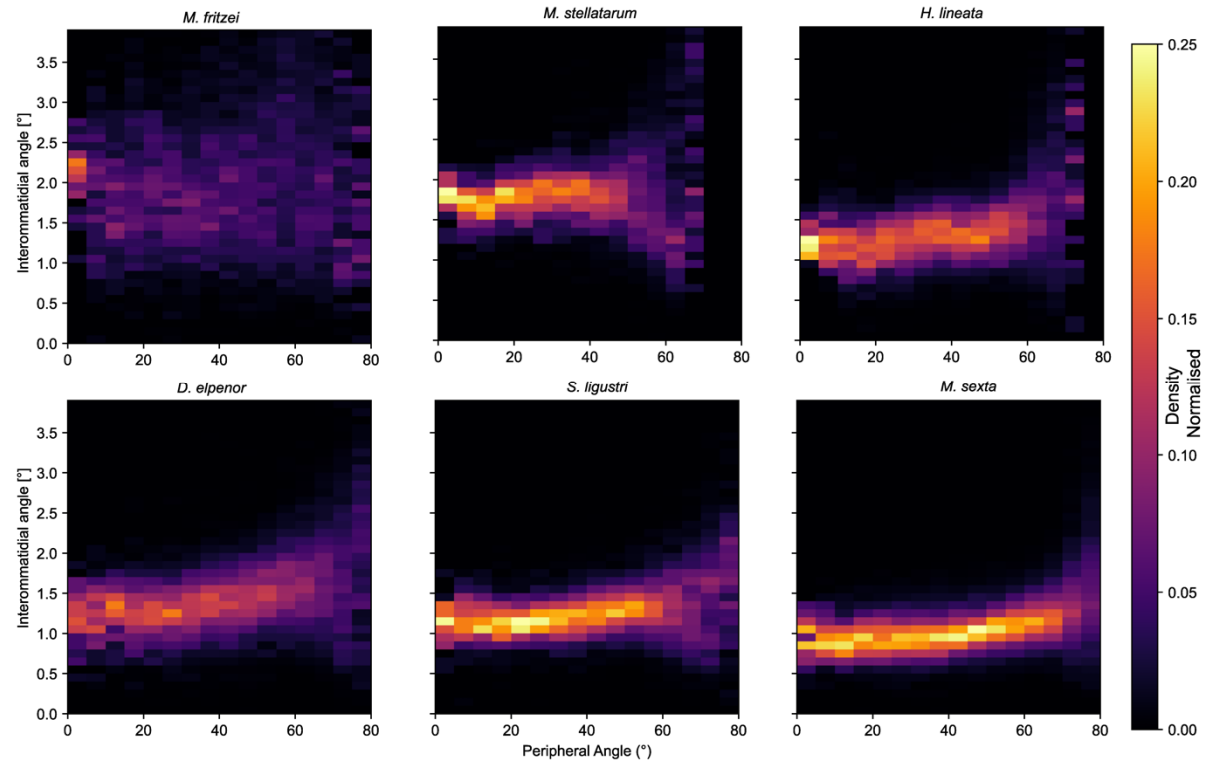


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Figure 2: Transverse μ CT scan cross-sections of the compound eye of the six hawkmoth species studied. Each cross-section passes through the center of the eye. In all images, the top corresponds to the anterior side of the head, and the bottom to the posterior.

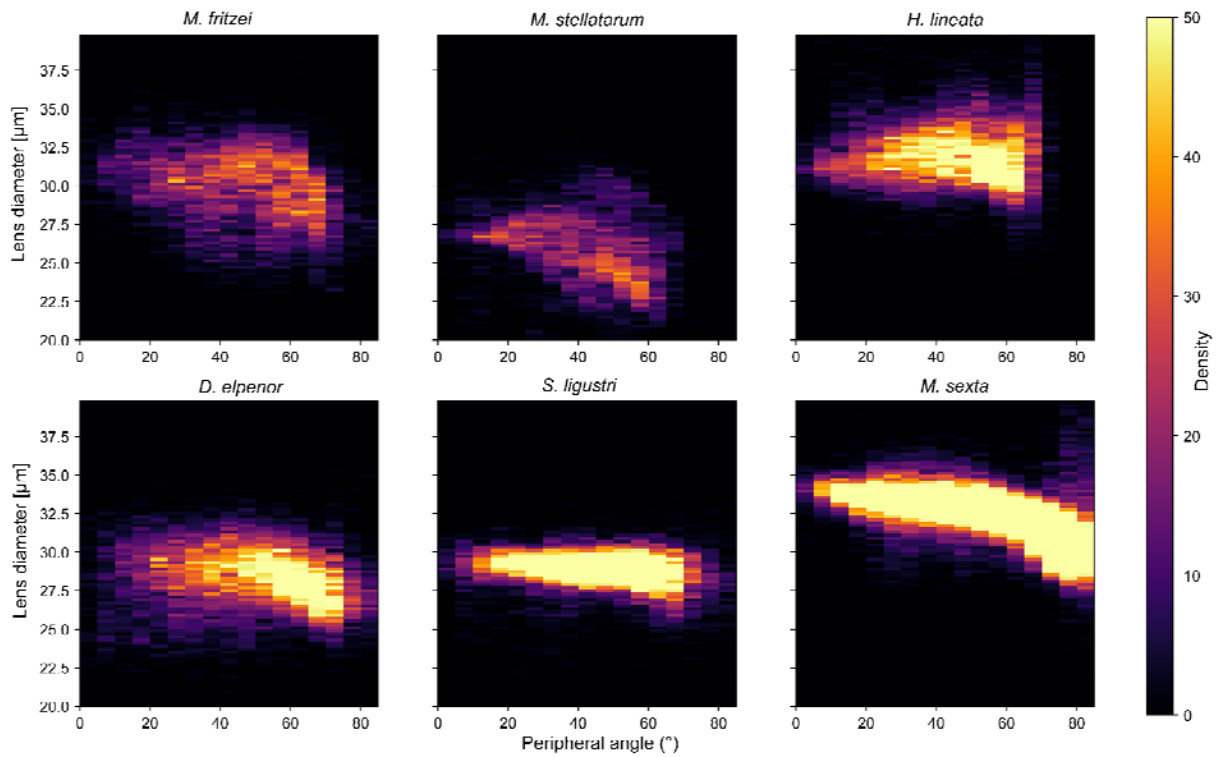


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524 **Figure 3:** Variation in estimated radius (sphericity) mapped onto the 3D surface of the eye to
525 showcase variation from center to dorsoventral (decrease) and anteroposterior edges (increase).
526 The histograms on the right indicate counts for the estimated radius values at various ommatidia.
527 Data were rotated to a global coordinate orientation.
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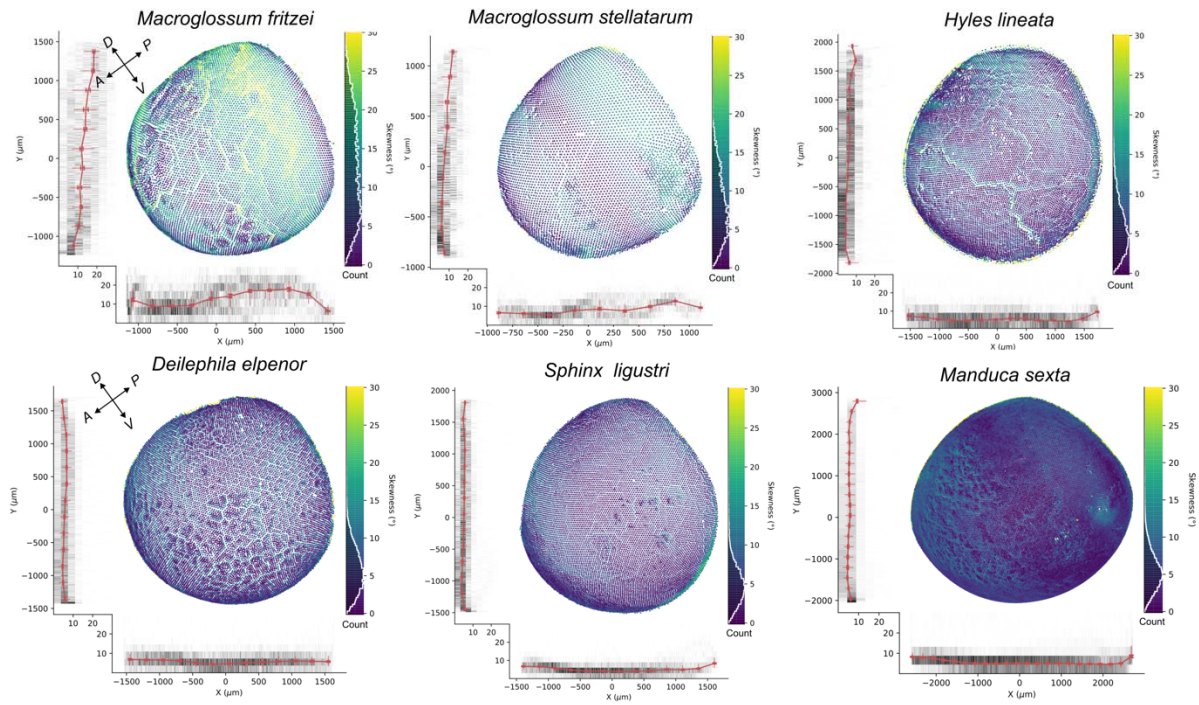


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Fig 4: Variation in $\Delta\phi$ from center to periphery. Species 1-3 (top row) have some diurnal activity and the three species on the bottom are nocturnal, axes and data has been truncated to showcase most variation and to make plots comparable.



535
536 **Fig 5:** Comparison of diel-activity change in relation to eye diameter and peripheral angle.
537 Peripheral angle was a measure of center or periphery, defined as $\frac{\theta + \phi}{2}$ where theta and
538 phi refer to the polar representations of the ommatidia. Species 1-3 (top row) exhibit diurnal
539 activity, and the three species on the bottom row are nocturnal. Species depicted (from upper left
540 to bottom right)
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Fig 6: Variation in cone skewness in X (posterior to anterior) and Y (ventral to dorsal) for all the hawkmoth species studied. Red lines depict individual counts along the X and Y axes. Data were rotated to a global coordinate orientation.

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Tables.

Table 1: List of species analyzed in this study. Parameters used for cleaning and the range of ODA output values for single eyes are listed. (D: Diurnal, N: Nocturnal)

Species	Diel Niche (D/N)	Specimen No.	Sex	Eye analyzed (R/L/Both)	CT Voxel (μm)	Facet counts	Facet diameter (μm)	$\Delta\phi$	Eye diameter (mm)
<i>Manduca sexta</i>	N	13460	-	Left	4.50074	28800	32.29 \pm 1.77	1.24 \pm 0.81	4.426
<i>Deilephila elpenor</i>	N	Bb_02	Male	Right	3.325	13068	28.21 \pm 1.86	1.55 \pm 0.75	2.754
<i>Hyles lineata</i>	D/N	HL002	Male	Right	3.4784	11922	32.05 \pm 1.85	1.41 \pm	2.967
<i>Macroglossum fritzei</i>	D/N	LEP29152	Male	Right	5.6845	8338	29.81 \pm 2.21	2.22 \pm 1.40	2.452
<i>Macroglossum stellatarum</i>	D	M002	Female	Right	2.4	5540	25.85 \pm 2.16	1.82 \pm 0.63	1.906
<i>Sphinx ligustri</i>	N	SLF02	Female	Right	3.3	12502	28.80 \pm 1.10	1.28 \pm 0.43	2.836
<i>Sphinx ligustri</i>	N	BB_09	Male	Both	4.7828	10955	NA	NA	2.937

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Supplemental Files

See link for supplementary data which includes figures, tables and protocols.

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Data Availability

Count data and ODA files are provided on

Figshare: <https://figshare.com/s/44be67162143d41fddb6>.

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The μCT scan for all six hawkmoth species analyzed in this study are publicly available on MorphoSource under the following DOI: 10.17602/M2/M780271; 10.17602/M2/M780277; 10.17602/M2/M780361; 10.17602/M2/M780378; 10.17602/M2/M780384; 10.17602/M2/M780405.

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